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W MASS

To obtain the world average, common systematics between experiments are properly taken into account. The LEP average W mass based on published results is 80.383 ± 0.035 GeV. The combined $p\bar{p}$ collider data yields an average W mass of 80.454 ± 0.059 GeV (ABAZOV 04D).

OUR FIT uses these average LEP and $p\overline{p}$ collider W mass values together with the Z mass, the W to Z mass ratio, and mass difference measurements.

VALUE (GeV)	EVTS	DOCUMENT ID		TECN	COMMENT
80.403± 0.029 OUR F	iT.				
$80.415 \pm 0.042 \pm 0.031$	11830	$^{ m 1}$ ABBIENDI	06	OPAL	E ^{ee} _{cm} = 170–209 GeV
$80.270 \pm 0.046 \pm 0.031$	9909	² ACHARD	06	L3	$E_{cm}^{ee} = 161-209 \text{ GeV}$
80.440± 0.043±0.027	8692	³ SCHAEL	06	ALEP	$E_{\mathrm{cm}}^{\mathrm{ee}} = 161 – 209 \; \mathrm{GeV}$
80.483 ± 0.084	49247	⁴ ABAZOV	02 D	D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
80.359± 0.074±0.049	3077	⁵ ABREU	01 K	DLPH	Eee = 161+172+183 +189 GeV
80.433± 0.079	53841	⁶ AFFOLDER	01E	CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
• • • We do not use t	he followir	ng data for averages	s, fits,	limits, e	
$82.87 \ \pm \ 1.82 \ ^{+0.30}_{-0.16}$	1500	⁷ AKTAS	06	H1	$e^{\pm} p ightarrow \overline{ u}_e(u_e) X$, $\sqrt{s} pprox 300 \; {\sf GeV}$
$80.3 \pm 2.1 \pm 1.2 \pm 1.0$	645	⁸ CHEKANOV	02 C	ZEUS	$e^- \stackrel{v}{p} \rightarrow \stackrel{v}{}_e X, \sqrt{s} = 318 \; GeV$
$81.4^{+2.7}_{-2.6} \pm 2.0^{+3.3}_{-3.0}$	1086	⁹ BREITWEG	00 D	ZEUS	$e^+p \rightarrow \overline{\nu}_e X, \sqrt{s} \approx 300 \text{ GeV}$
$80.84 \pm 0.22 \pm 0.83$	2065	¹⁰ ALITTI	92B	UA2	See W/Z ratio below
$80.79 \pm 0.31 \pm 0.84$		¹¹ ALITTI	90 B	UA2	$E_{ m cm}^{p\overline{p}}=$ 546,630 GeV
$80.0 \pm 3.3 \pm 2.4$	22	¹² ABE	891	CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
82.7 \pm 1.0 \pm 2.7	149	¹³ ALBAJAR	89	UA1	$E_{ m cm}^{p\overline{p}}=$ 546,630 GeV
81.8 $\begin{array}{ccc} + & 6.0 \\ - & 5.3 \end{array} \pm 2.6$	46	¹⁴ ALBAJAR	89	UA1	$E_{\rm cm}^{p\overline{p}}$ = 546,630 GeV
$89 \pm 3 \pm 6$	32	¹⁵ ALBAJAR	89	UA1	$E_{ m cm}^{p\overline{p}}=$ 546,630 GeV
81. ± 5.	6	ARNISON	83	UA1	Eee 546 GeV
80. $ +10. $ $-6. $	4	BANNER	83 B	UA2	Repl. by ALITTI 90B

¹ ABBIENDI 06 use direct reconstruction of the kinematics of $W^+W^- \to q \overline{q} \ell \nu_\ell$ and $W^+W^- \to q \overline{q} q \overline{q}$ events. The result quoted here is obtained combining this mass value with the results using $W^+W^- \to \ell \nu_\ell \ell' \nu_{\ell'}$ events in the energy range 183–207 GeV (ABBIENDI 03C) and the dependence of the WW production cross-section on m_W at threshold. The systematic error includes ± 0.009 GeV due to the uncertainty on the LEP beam energy.

- ² ACHARD 06 use direct reconstruction of the kinematics of $W^+W^- \to q \overline{q} \ell \nu_\ell$ and $W^+W^- \to q \overline{q} q \overline{q}$ events in the C.M. energy range 189–209 GeV. The result quoted here is obtained combining this mass value with the results obtained from a direct W mass reconstruction at 172 and 183 GeV and with those from the dependence of the WW production cross-section on m_W at 161 and 172 GeV (ACCIARRI 99).
- ³ SCHAEL 06 use direct reconstruction of the kinematics of $W^+W^- \to q \overline{q} \ell \nu_\ell$ and $W^+W^- \to q \overline{q} q \overline{q}$ events in the C.M. energy range 183–209 GeV. The result quoted here is obtained combining this mass value with those obtained from the dependence of the W pair production cross-section on m_W at 161 and 172 GeV (BARATE 97 and BARATE 97S respectively). The systematic error includes ± 0.009 GeV due to possible effects of final state interactions in the $q \overline{q} q \overline{q}$ channel and ± 0.009 GeV due to the uncertainty on the LEP beam energy.
- ⁴ABAZOV 02D improve the measurement of the W-boson mass including $W \to e \nu_e$ events in which the electron is close to a boundary of a central electromagnetic calorimeter module. Properly combining the results obtained by fitting $m_T(W)$, $p_T(e)$, and $p_T(\nu)$, this sample provides a mass value of 80.574 \pm 0.405 GeV. The value reported here is a combination of this measurement with all previous DØ W-boson mass measurements.
- ⁵ ABREU 01K obtain this value properly combining results obtained from a direct W mass reconstruction at 172, 183, and 189 GeV with those from measurements of W-pair production cross sections at 161, 172, and 183 GeV. The systematic error includes ± 0.017 GeV due to the beam energy uncertainty and ± 0.033 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.
- 6 AFFOLDER 01E fit the transverse mass spectrum of 30115 $W\to e\nu_e$ events ($M_W=80.473\pm0.065\pm0.092$ GeV) and of 14740 $W\to \mu\nu_\mu$ events ($M_W=80.465\pm0.100\pm0.103$ GeV) obtained in the run IB (1994-95). Combining the electron and muon results, accounting for correlated uncertainties, yields $M_W=80.470\pm0.089$ GeV. They combine this value with their measurement of ABE 95P reported in run IA (1992-93) to obtain the quoted value.
- 7 AKTAS 06 fit the Q^2 dependence (300 < Q^2 < 30,000 $\mbox{GeV}^2)$ of the charged-current differential cross section with a propagator mass. The first error is experimental and the second corresponds to uncertainties due to input parameters and model assumptions.
- ⁸ CHEKANOV 02C fit the Q^2 dependence (200< Q^2 <60000 GeV 2) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.
- ⁹ BREITWEG 00D fit the Q^2 dependence (200 < Q^2 < 22500 GeV²) of the charged-current differential cross sections with a propagator mass fit. The last error is due to the uncertainty on the probability density functions.
- 10 ALITTI 92B result has two contributions to the systematic error (± 0.83); one (± 0.81) cancels in m_W/m_Z and one (± 0.17) is noncancelling. These were added in quadrature. We choose the ALITTI 92B value without using the LEP m_Z value, because we perform our own combined fit.
- ¹¹ There are two contributions to the systematic error (± 0.84): one (± 0.81) which cancels in m_W/m_Z and one (± 0.21) which is non-cancelling. These were added in quadrature.
- 12 ABE 891 systematic error dominated by the uncertainty in the absolute energy scale.
- 13 ALBAJAR 89 result is from a total sample of 299 W
 ightarrow ~e
 u events.
- 14 ALBAJAR 89 result is from a total sample of 67 $W
 ightarrow ~\mu
 u$ events.
- 15 ALBAJAR 89 result is from W
 ightarrow au
 u events.

W/Z MASS RATIO

The fit uses the W and Z mass, mass difference, and mass ratio measurements.

VALUE	EVTS	DOCUMENT ID		TECN	COMMENT
0.88173 ± 0.00032 OUR FIT					
$0.8821\ \pm0.0011\ \pm0.0008$	28323	¹⁶ ABBOTT	98N	D0	$E_{cm}^{oldsymbol{p}\overline{oldsymbol{p}}}$ $= 1.8\;TeV$
$0.88114 \pm 0.00154 \pm 0.00252$	5982	¹⁷ ABBOTT	98 P	D0	$E_{cm}^{oldsymbol{p}\overline{oldsymbol{p}}}$ $= 1.8\;TeV$
$0.8813 \pm 0.0036 \pm 0.0019$	156	¹⁸ ALITTI	92 B	UA2	$E_{\rm cm}^{p\overline{p}}=$ 630 GeV

 $^{^{16}}$ ABBOTT 98N obtain this from a study of 28323 $W \to e \nu_e$ and 3294 $Z \to e^+ e^-$ decays. Of this latter sample, 2179 events are used to calibrate the electron energy scale.

$m_Z - m_W$

The fit uses the W and Z mass, mass difference, and mass ratio measurements.

VALUE	(GeV)		DOCUMENT ID		TECN	COMMENT
10.78	5 ± 0.02	9 OUR FIT				
10.4	±1.4	±0.8	ALBAJAR	89	UA1	$E_{cm}^{p\overline{p}} = 546,630 \; GeV$
• • •	We do	not use the followi	ng data for average	es, fits,	limits,	etc. • • •
11.3	± 1.3	±0.9	ANSARI	87	UA2	$E_{\rm cm}^{p \overline{p}} = 546,630 \; {\rm GeV}$

$m_{W^+} - m_{W^-}$

Test of CPT invariance.

VALUE (GeV)	EVTS	DOCUMENT ID		TECN	COMMENT
-0.19 ± 0.58	1722	ABE	90G	CDF	$E_{cm}^{oldsymbol{\overline{p}}} = 1.8 \; TeV$

W WIDTH

The CDF and DØ widths labelled "extracted value" are obtained by measuring $R = [\sigma(W)/\sigma(Z)] \ [\Gamma(W \to \ell \nu_\ell)]/(\mathrm{B}(Z \to \ell \ell)\Gamma(W))$ where the bracketed quantities can be calculated with plausible reliability. $\Gamma(W)$ is then extracted by using a value of $\mathrm{B}(Z \to \ell \ell)$ measured at LEP. The UA1 and UA2 widths used $R = [\sigma(W)/\sigma(Z)] \ [\Gamma(W \to \ell \nu_\ell)/\Gamma(Z \to \ell \ell)] \ \Gamma(Z)/\Gamma(W)$ and the measured value of $\Gamma(Z)$. The Standard Model prediction is 2.0910 ± 0.0015 GeV (see Review on "Electroweak model and constraints on new physics" in this Edition).

To obtain OUR FIT, the correlation between systematics within LEP experiments and within Tevatron experiments is properly taken into account as given in the LEP note accessible at http://lepewwg.web.cern.ch/LEPEWWG/lepww/mw/pdg_2006/ and in the combined Tevatron paper of ABAZOV 04D. The respective average

 $^{^{17}}$ ABBOTT 98P obtain this from a study of 5982 $W\to e\nu_e$ events. The systematic error includes an uncertainty of ± 0.00175 due to the electron energy scale.

¹⁸ Scale error cancels in this ratio.

values (2.164 \pm 0.085 GeV from LEP and 2.115 \pm 0.105 GeV from Tevatron) yield an average W width of 2.145 \pm 0.066 GeV coming from direct measurements. ABAZOV 04D also determine the average extracted W width using CDF and DØ data to obtain a value of 2.141 \pm 0.057 GeV.

They further combine the Tevatron direct and extracted W widths to obtain an average Tevatron width of 2.135 ± 0.050 GeV. Finally combining this with the LEP W width and the extracted W width values from UA1 and UA2 one obtains the quoted value.

VALUE (GeV)	CL% EVTS	DOCUMENT ID)	TECN	COMMENT
2.141±0.041 OUR	FIT				
$1.996 \pm 0.096 \pm 0.10$	2 10729	¹⁹ ABBIENDI	06	OPAL	E ^{ee} _{cm} = 170–209 GeV
$2.18 \pm 0.11 \pm 0.09$	9795	²⁰ ACHARD	06	L3	E ^{ee} _{cm} = 172–209 GeV
$2.14 \pm 0.09 \pm 0.06$	8717	²¹ SCHAEL	06	ALEP	$E_{\rm cm}^{\it ee} = 183 – 209 \; {\rm GeV}$
$2.23 \ ^{+0.15}_{-0.14} \ \pm 0.10$	294	²² ABAZOV	02E	D0	Direct meas.
$2.266 \pm 0.176 \pm 0.07$	6 3005	²³ ABREU	01K	DLPH	$E_{\rm cm}^{\it ee} = 183,189 \; {\rm GeV}$
$2.152\!\pm\!0.066$	79176	²⁴ ABBOTT	00 B	D0	Extracted value
$2.05 \pm 0.10 \pm 0.08$	662	²⁵ AFFOLDER	00M	CDF	Direct meas.
$2.064 \pm 0.060 \pm 0.05$	9	²⁶ ABE	95W	CDF	Extracted value
$2.10 \ ^{+0.14}_{-0.13} \ \pm 0.09$	3559	²⁷ ALITTI	92	UA2	Extracted value
$2.18 \ ^{+0.26}_{-0.24} \ \pm 0.04$	ŀ	²⁸ ALBAJAR	91	UA1	Extracted value
• • • We do not us	se the followin	g data for average	es, fit	s, limits	, etc. • • •
$2.30 \pm 0.19 \pm 0.06$		²⁹ ALITTI	90 C	UA2	Extracted value
$2.8 \begin{array}{cc} +1.4 \\ -1.5 \end{array} \pm 1.3$	149	³⁰ ALBAJAR	89	UA1	$E_{cm}^{p\overline{p}} = 546,630 \; GeV$
< 7	90 119	APPEL	86	UA2	$E_{\rm cm}^{p\overline{p}} = 546,630 \; {\rm GeV}$
< 6.5	90 86	³¹ ARNISON	86	UA1	$E_{\rm cm}^{p\overline{p}}$ = 546,630 GeV

¹⁹ ABBIENDI 06 use direct reconstruction of the kinematics of $W^+W^-\to q\overline{q}\ell\nu_\ell$ and $W^+W^-\to q\overline{q}q\overline{q}$ events. The systematic error includes ± 0.003 GeV due to the uncertainty on the LEP beam energy.

²⁰ ACHARD 06 use direct reconstruction of the kinematics of $W^+W^- \to q \overline{q} \ell \nu_\ell$ and $W^+W^- \to q \overline{q} q \overline{q}$ events in the C.M. energy range 189–209 GeV. The result quoted here is obtained combining this value of the width with the result obtained from a direct W mass reconstruction at 172 and 183 GeV (ACCIARRI 99).

²¹ SCHAEL 06 use direct reconstruction of the kinematics of $W^+W^- \to q \overline{q} \ell \nu_\ell$ and $W^+W^- \to q \overline{q} q \overline{q}$ events. The systematic error includes ± 0.05 GeV due to possible effects of final state interactions in the $q \overline{q} q \overline{q}$ channel and ± 0.01 GeV due to the uncertainty on the LEP beam energy.

²² ABAZOV 02E obtain this result fitting the high-end tail (90–200 GeV) of the transverse-mass spectrum in semileptonic $W\to e\nu_e$ decays.

²³ ABREU 01K obtain this value properly combining results obtained at 183 and 189 GeV using $WW \to \ell \overline{\nu}_\ell q \overline{q}$ and $WW \to q \overline{q} q \overline{q}$ decays. The systematic error includes an uncertainty of ± 0.052 GeV due to possible color reconnection and Bose-Einstein effects in the purely hadronic final state.

ABBOTT 00B measure $R=10.43\pm0.27$ for the $W\to e\nu_e$ decay channel. They use the SM theoretical predictions for $\sigma(W)/\sigma(Z)$ and $\Gamma(W\to e\nu_e)$ and the world average for B($Z\to ee$). The value quoted here is obtained combining this result (2.169 \pm 0.070 GeV) with that of ABBOTT 99H.

- $^{25}\, {\rm AFFOLDER}$ 00M fit the high transverse mass (100–200 GeV) $W\to e\nu_e$ and $W\to \mu\nu_\mu$ events to obtain $\Gamma(W)=2.04\pm0.11({\rm stat})\pm0.09({\rm syst})$ GeV. This is combined with the earlier CDF measurement (ABE 95C) to obtain the quoted result.
- ²⁶ ABE 95W measured $R=10.90\pm0.32\pm0.29$. They use $m_{W}=80.23\pm0.18$ GeV, $\sigma(W)/\sigma(Z)=3.35\pm0.03$, $\Gamma(W\to e\nu)=225.9\pm0.9$ MeV, $\Gamma(Z\to e^+e^-)=83.98\pm0.18$ MeV, and $\Gamma(Z)=2.4969\pm0.0038$ GeV.
- ²⁷ ALITTI 92 measured $R=10.4^{+0.7}_{-0.6}\pm0.3$. The values of $\sigma(Z)$ and $\sigma(W)$ come from $O(\alpha_s^2)$ calculations using $m_W=80.14\pm0.27$ GeV, and $m_Z=91.175\pm0.021$ GeV along with the corresponding value of $\sin^2\theta_W=0.2274$. They use $\sigma(W)/\sigma(Z)=3.26\pm0.07\pm0.05$ and $\Gamma(Z)=2.487\pm0.010$ GeV.
- 3.26 \pm 0.07 \pm 0.05 and $\Gamma(Z) = 2.487 \pm$ 0.010 GeV. ²⁸ ALBAJAR 91 measured $R = 9.5 ^{+1.1}_{-1.0}$ (stat. + syst.). $\sigma(W)/\sigma(Z)$ is calculated in QCD at the parton level using $m_W = 80.18 \pm 0.28$ GeV and $m_Z = 91.172 \pm 0.031$ GeV along with $\sin^2 \theta_W = 0.2322 \pm 0.0014$. They use $\sigma(W)/\sigma(Z) = 3.23 \pm 0.05$ and $\Gamma(Z) = 2.498 \pm 0.020$ GeV. This measurement is obtained combining both the electron and muon channels
- muon channels. 29 ALITTI 90C used the same technique as described for ABE 90. They measured $R=9.38^{+0.82}_{-0.72}\pm0.25$, obtained $\Gamma(W)/\Gamma(Z)=0.902\pm0.074\pm0.024$. Using $\Gamma(Z)=2.546\pm0.032$ GeV, they obtained the $\Gamma(W)$ value quoted above and the limits $\Gamma(W)<<2.56$ (2.64) GeV at the 90% (95%) CL. $E_{\rm CM}^{P\overline{p}}=546,630$ GeV.
- 30 ALBAJAR 89 result is from a total sample of 299 W
 ightarrow e
 u events.
- 31 If systematic error is neglected, result is $2.7^{+1.4}_{-1.5}$ GeV. This is enhanced subsample of 172 total events.

W⁺ DECAY MODES

 W^- modes are charge conjugates of the modes below.

	Mode	Fraction (Γ_i/Γ)	Confidence level
$\overline{\Gamma_1}$	$\ell^+ \nu$	[a] (10.80± 0.09) %	_
	$e^+ \nu$	$(10.75 \pm 0.13) \%$	
	$\mu^+ \nu$	$(10.57 \pm \ 0.15) \%$	
Γ_4	$ au^+ u$	$(11.25 \pm 0.20) \%$	
Γ_5	hadrons	$(67.60 \pm 0.27) \%$	
	$\pi^+ \gamma$	< 8 × 1	10^{-5} 95%
Γ_7	$D_s^+ \gamma$	< 1.3 × 1	10^{-3} 95%
Γ ₈	cX	(33.4 \pm 2.6) %	
Γ ₉	C S	$(31 {}^{+13}_{-11}) \%$	
Γ_{10}	invisible	[b] (1.4 \pm 2.8) %	

- [a] ℓ indicates each type of lepton $(e, \mu, \text{ and } \tau)$, not sum over them.
- [b] This represents the width for the decay of the W boson into a charged particle with momentum below detectability, p< 200 MeV.

W PARTIAL WIDTHS

 Γ (invisible) Γ_{10}

This represents the width for the decay of the W boson into a charged particle with momentum below detectability, p< 200 MeV.

VALUE (MeV)DOCUMENT IDTECNCOMMENT $30^{+52}_{-48} \pm 33$ 32 BARATE99IALEP $E^{ee}_{cm} = 161 + 172 + 183$
GeV

• • We do not use the following data for averages, fits, limits, etc. • •

33 BARATE 99L ALEP $E_{cm}^{ee} = 161 + 172 + 183$ GeV

W BRANCHING RATIOS

Overall fits are performed to determine the branching ratios of the W. LEP averages on $W \rightarrow e \nu_e$, $W \rightarrow \mu \nu_\mu$, and $W \rightarrow \tau \nu_\tau$, and their correlations are first obtained by combining results from the four experiments taking properly into account the common systematics. The procedure is described in the note LEPEWWG/XSEC/2001-02, 30 March 2001, at http://lepewwg.web.cern.ch/LEPEWWG/lepww/4f/PDG01. The LEP average values so obtained, using published data, are given in the note LEPEWWG/XSEC/2005-01 accessible at http://lepewwg.web.cern.ch/ LEPEWWG/lepww/4f/PDG05/. These results, together with results from the $p\bar{p}$ colliders are then used in fits to obtain the world average W branching ratios. A first fit determines three individual leptonic branching ratios, $\mathsf{B}(W \to e \nu_e)$, $\mathsf{B}(W \to \mu \nu_\mu)$, and $\mathsf{B}(W \to \tau \nu_\tau)$. This fit has a $\chi^2 =$ 4.7 for 10 degrees of freedom. A second fit assumes lepton universality and determines the leptonic branching ratio B($W \rightarrow \ell \nu_{\ell}$) and the hadronic branching ratio is derived as B($W \to \text{hadrons}$) = 1–3 B($W \to \ell \nu$). This fit has a $\chi^2=11.3$ for 12 degrees of freedom.

The LEP $W \to \ell \nu$ data are obtained by the Collaborations using individual leptonic channels and are, therefore, not included in the overall fits to avoid double counting.

 $\Gamma(\ell^+
u)/\Gamma_{ ext{total}}$ ℓ indicates average over e, μ , and au modes, not sum over modes.

<u>VALUE</u>	EVTS	DOCUMENT ID		TECN	COMMENT
0.1080±0.0009 OUR FIT	•				
$0.1085\!\pm\!0.0014\!\pm\!0.0008$	13600	ABDALLAH			$E_{\rm cm}^{\it ee} = 161 - 209 \; {\rm GeV}$
$0.1083\!\pm\!0.0014\!\pm\!0.0010$	11246	ACHARD	04 J	L3	$E_{\rm cm}^{\rm ee} = 161 - 209 \; {\rm GeV}$
$0.1096\!\pm\!0.0012\!\pm\!0.0005$	16116	SCHAEL	04A	ALEP	$E_{\rm cm}^{\it ee} = 183-209 \; {\rm GeV}$
$0.1056\!\pm\!0.0020\!\pm\!0.0009$	5778	ABBIENDI,G	00	OPAL	$E_{\rm cm}^{\rm ee} = 161 + 172 + 183$
					<u>+</u> 189 GeV
$0.1102\!\pm\!0.0052$	11858	³⁴ ABBOTT	99н	D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
0.104 ± 0.008	3642	³⁵ ABE	921	CDF	$E_{cm}^{p\overline{\overline{p}}} = 1.8 \; TeV$
LITTO //DDC.LDL.C	O) /	Б 6		_	

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 $^{^{32}\,\}text{BARATE}$ 991 measure this quantity using the dependence of the total cross section $\sigma_{W\,W}$ upon a change in the total width. The fit is performed to the $W\,W$ measured cross sections at 161, 172, and 183 GeV. This partial width is < 139 MeV at 95%CL.

³³ BARATE 99L use W-pair production to search for effectively invisible W decays, tagging with the decay of the other W boson to Standard Model particles. The partial width for effectively invisible decay is < 27 MeV at 95%CL.

³⁴ ABBOTT 99H measure $R \equiv [\sigma_W \ \mathsf{B}(W \to \ell \nu_\ell)]/[\sigma_Z \ \mathsf{B}(Z \to \ell \ell)] = 10.90 \pm 0.52$ combining electron and muon channels. They use $M_W = 80.39 \pm 0.06$ GeV and the SM theoretical predictions for $\sigma(W)/\sigma(Z)$ and $\mathsf{B}(Z \to \ell \ell)$.

³⁵ $1216 \pm 38^{+27}_{-31} \ W \to \mu \nu$ events from ABE 92I and 2426 $W \to e \nu$ events of ABE 91C.

ABE 921 give the inverse quantity as 9.6 \pm 0.7 and we have inverted.

$\Gamma(e^+ u)/\Gamma_{ m total}$					Γ_2/Γ
VALUE	EVTS	DOCUMENT ID		TECN	COMMENT
0.1075±0.0013 OUR FIT					<u></u>
0.1061 ± 0.0028		³⁶ ABAZOV	04 D	TEVA	$E_{cm}^{ar{p}} = 1.8 \; TeV$
$0.1055 \!\pm\! 0.0031 \!\pm\! 0.0014$	1804	ABDALLAH	04 G	DLPH	$E_{\rm cm}^{\it ee} = 161 - 209 \; {\rm GeV}$
$0.1078\!\pm\!0.0029\!\pm\!0.0013$	1576	ACHARD	04 J	L3	$E_{\rm cm}^{\it ee} = 161 - 209 \; {\rm GeV}$
$0.1078\!\pm\!0.0027\!\pm\!0.0010$	2142	SCHAEL	04A	ALEP	$E_{\rm cm}^{\it ee} = 183 – 209 \; {\rm GeV}$
$0.1046 \pm 0.0042 \pm 0.0014$	801	ABBIENDI,G	00	OPAL	E ^{ee} _{cm} = 161+172+183 +189 GeV
0.10 ± 0.014 $^{+0.02}_{-0.03}$	248	³⁷ ANSARI	87 C	UA2	$E_{cm}^{p\overline{p}} = 546,630 \; GeV$
• • • We do not use the f	following	data for averages,	fits, li	mits, etc	c. • • •
seen	119	APPEL	86	UA2	$E_{cm}^{p\overline{p}} = 546,630 \; GeV$
seen	172	ARNISON	86	UA1	$E_{\rm cm}^{p\overline{p}} = 546,630 \; {\rm GeV}$

 36 ABAZOV 04D take into account all correlations to properly combine the CDF (ABE 95W) and DØ (ABBOTT 00B) measurements of the ratio R in the electron channel. The ratio R is defined as $[\sigma_W \cdot B(W \to e \nu_e)] / [\sigma_Z \cdot B(Z \to ee)]$. The combination gives R $^{Tevatron}=$ 10.59 \pm 0.23. σ_W / σ_Z is calculated at next-to-next-to-leading order (3.360 \pm 0.051). The branching fraction B($Z \rightarrow ee$) is taken from this Reviewas $(3.363 \pm 0.004)\%$.

37 The first error was obtained by adding the statistical and systematic experimental uncertainties in quadrature. The second error reflects the dependence on theoretical prediction of total W cross section: $\sigma(546~{\rm GeV})=4.7^{+1.4}_{-0.7}$ nb and $\sigma(630~{\rm GeV})=5.8^{+1.8}_{-1.0}$ nb. See ALTARELLI 85B.

$\Gamma(\mu^+ u)/\Gamma_{total}$					Г3/Г
VALUE	EVTS	DOCUMENT ID		TECN	COMMENT
0.1057±0.0015 OUR FIT					
$0.1065\!\pm\!0.0026\!\pm\!0.0008$	1998	ABDALLAH	04 G	DLPH	$E_{\rm cm}^{\rm ee} = 161 - 209 \; {\rm GeV}$
$0.1003\!\pm\!0.0029\!\pm\!0.0012$	1423	ACHARD	04 J	L3	$E_{\mathrm{cm}}^{\mathrm{ee}} = 161209 \; \mathrm{GeV}$
$0.1087 \!\pm\! 0.0025 \!\pm\! 0.0008$	2216	SCHAEL	04A	ALEP	$E_{\rm cm}^{\it ee} = 183-209 \; {\rm GeV}$
$0.1050 \pm 0.0041 \pm 0.0012$	803	ABBIENDI,G	00	OPAL	$E_{\rm cm}^{\rm ee} =$
					161+172+183 <u>+</u> 189 GeV
0.10 ± 0.01	1216	³⁸ ABE	921	CDF	$_{cm}^{+189} GeV$ $E_{cm}^{oldsymbol{p}} = 1.8 TeV$

 $^{^{38}}$ ABE 921 quote the inverse quantity as 9.9 \pm 1.2 which we have inverted.

$\Gamma(\tau^+ u)/\Gamma_{ m total}$					Γ_4/Γ
VALUE	<u>EVTS</u>	DOCUMENT ID		TECN	COMMENT
0.1125±0.0020 OUR FIT					
$0.1146\!\pm\!0.0039\!\pm\!0.0019$	2034	ABDALLAH	04 G	DLPH	$E_{cm}^{ee} = 161209 \; GeV$
$0.1189\!\pm\!0.0040\!\pm\!0.0020$	1375	ACHARD	04 J	L3	$E_{cm}^{ee} = 161209 \; GeV$
$0.1125\!\pm\!0.0032\!\pm\!0.0020$	2070	SCHAEL	04A	ALEP	$E_{\rm cm}^{\it ee} = 183-209 \; {\rm GeV}$
$0.1075 \pm 0.0052 \pm 0.0021$	794	ABBIENDI,G	00	OPAL	E ^{ee} _{Cm} = 161+172+183 +189 GeV

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 $\Gamma(\text{hadrons})/\Gamma_{\text{total}}$

OUR FIT value is obtained by a fit to the lepton branching ratio data assuming lepton universality.

 Γ_5/Γ

VALUE	EVTS	DOCUMENT ID		TECN	COMMENT
0.6760±0.0027 OUR FIT	•				
$0.6745 \pm 0.0041 \pm 0.0024$	13600	ABDALLAH	04 G	DLPH	$E_{cm}^{ee} = 161209 \; GeV$
$0.6750 \pm 0.0042 \pm 0.0030$	11246	ACHARD	04 J	L3	$E_{cm}^{ee} = 161209 \; GeV$
$0.6713 \!\pm\! 0.0037 \!\pm\! 0.0015$	16116	SCHAEL	04A	ALEP	$E_{cm}^{ee} = 183209 \; GeV$
$0.6832 \!\pm\! 0.0061 \!\pm\! 0.0028$	5778	ABBIENDI,G	00	OPAL	$E_{\rm cm}^{\it ee}=161+172+183$
					$+189 \; GeV$

 $\Gamma(\mu^+\nu)/\Gamma(e^+\nu)$ Γ_3/Γ_2 TECN COMMENT **EVTS** 0.983 ± 0.018 OUR FIT $E_{\rm cm}^{p\overline{p}} = 1.8 \text{ TeV}$ ³⁹ ABACHI 95D D0 0.89 ± 0.10 13k $E_{\rm cm}^{p\overline{p}} = 1.8 \text{ TeV}$ ⁴⁰ ABE CDF 1216 1.02 ± 0.08 E^{pp}= 546,630 GeV $1.00 \pm 0.14 \pm 0.08$ 67 **ALBAJAR** UA1

ullet ullet We do not use the following data for averages, fits, limits, etc. ullet ullet

$$1.24 \buildrel {+0.6}{-0.4} \buildrel {-0.4}$$
 14 ARNISON 84D UA1 Repl. by ALBA-JAR 89

⁴⁰ ABE 92I obtain σ_W B($W \to \mu\nu$)= 2.21 \pm 0.07 \pm 0.21 and combine with ABE 91C σ_W B(($W \to e\nu$)) to give a ratio of the couplings from which we derive this measurement.

$\Gamma(au^+ u)/\Gamma(e^+ u)$					Γ_4/Γ_2
VALUE	EVTS	DOCUMENT ID		TECN	COMMENT
1.046 ± 0.023 OUR FIT					
0.961 ± 0.061	980	⁴¹ ABBOTT	00 D	D0	$E_{CM}^{ar{p}} = 1.8 \; TeV$
$0.94\ \pm0.14$	179	⁴² ABE	92E	CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$
$1.04\ \pm0.08\ \pm0.08$	754	⁴³ ALITTI	92F	UA2	$E_{ m cm}^{p \overline{p}} =$ 630 GeV
$1.02 \ \pm 0.20 \ \pm 0.12$	32	ALBAJAR	89	UA1	$E_{\mathrm{cm}}^{p\overline{p}}$ = 546,630 GeV
• • • We do not use the	following	data for averages,	fits, li	mits, et	C. ● ● ●
$0.995\!\pm\!0.112\!\pm\!0.083$	198	ALITTI	91 C	UA2	Repl. by ALITTI 92F
$1.02 \pm 0.20 \pm 0.10$	32	ALBAJAR	87	UA1	Repl. by ALBA- JAR 89

⁴¹ ABBOTT 00D measure $\sigma_W \times {\sf B}(W \to \tau \nu_\tau) = 2.22 \pm 0.09 \pm 0.10 \pm 0.10$ nb. Using the ABBOTT 00B result $\sigma_W \times {\sf B}(W \to e \nu_e) = 2.31 \pm 0.01 \pm 0.05 \pm 0.10$ nb, they quote the ratio of the couplings from which we derive this measurement.

³⁹ ABACHI 95D obtain this result from the measured σ_W B($W \rightarrow \mu \nu$)= 2.09 \pm 0.23 \pm 0.11 nb and σ_W B($W \rightarrow e \nu$)= 2.36 \pm 0.07 \pm 0.13 nb in which the first error is the combined statistical and systematic uncertainty, the second reflects the uncertainty in the luminosity.

⁴² ABE 92E use two procedures for selecting $W \to \tau \nu_{\tau}$ events. The missing E_T trigger leads to $132 \pm 14 \pm 8$ events and the τ trigger to $47 \pm 9 \pm 4$ events. Proper statistical and systematic correlations are taken into account to arrive at $\sigma B(W \to \tau \nu) = 2.05 \pm 0.27$ nb. Combined with ABE 91C result on $\sigma B(W \to e\nu)$, ABE 92E quote a ratio of the couplings from which we derive this measurement.

 $^{^{43}}$ This measurement is derived by us from the ratio of the couplings of ALITTI 92F.

$\Gamma(\pi^+\gamma)/\Gamma(e^+ u)$					Γ_6/Γ_2
VALUE	CL%	DOCUMENT ID		TECN	COMMENT
$< 7 \times 10^{-4}$	95	ABE	98н	CDF	$E_{cm}^{ar{p}} = 1.8 \; TeV$
$< 4.9 \times 10^{-3}$	95	⁴⁴ ALITTI	92 D	UA2	$E_{ m cm}^{{ar p}} =$ 630 GeV
$< 58 \times 10^{-3}$	95	⁴⁵ ALBAJAR	90	UA1	$E_{ m cm}^{p\overline{p}} =$ 546, 630 GeV
44 ALITTI 92D limit i 45 ALBA IAR 90 obta	s 3.8×10	0 ⁻³ at 90%CL.			

$\Gamma(D_s^+\gamma)/\Gamma(e^+ u)$						Γ_7/Γ_2
VALUE	CL%	DOCUMENT ID		TECN	COMMENT	
$<1.2 \times 10^{-2}$	95	ABE	98 P	CDF	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$	
$\Gamma(cX)/\Gamma(hadrons)$						Γ_8/Γ_5
VALUE	EVTS	DOCUMENT ID		TECN	COMMENT	
0.49 ±0.04 OUR AVE	RAGE					
$0.481\!\pm\!0.042\!\pm\!0.032$	3005	⁴⁶ ABBIENDI	00V	OPAL	$E_{\rm cm}^{\it ee} = 183 + 18$	9 GeV
$0.51 \ \pm 0.05 \ \pm 0.03$	746	⁴⁷ BARATE	99м	ALEP	$E_{\rm cm}^{\it ee} = 172 + 18$	3 GeV

 $^{^{46}}$ ABBIENDI 00V tag W
ightarrow c X decays using measured jet properties, lifetime information, and leptons produced in charm decays. From this result, and using the additional measurements of $\Gamma(W)$ and $B(W \to hadrons)$, $|V_{CS}|$ is determined to be $0.969 \pm 0.045 \pm 0.036$.

47 BARATE 99M tag c jets using a neural network algorithm. From this measurement $|V_{cs}|$ is determined to be 1.00 \pm 0.11 \pm 0.07.

$R_{cs} = \Gamma(c\overline{s})/\Gamma(hadrons)$				Γ_9/Γ_5
VALUE	DOCUMENT ID		TECN	COMMENT
$0.46^{+0.18}_{-0.14}\pm0.07$	⁴⁸ ABREU	98N	DLPH	E ^{ee} _{cm} = 161+172 GeV

 $^{^{}m 48}$ ABREU 98N tag c and s jets by identifying a charged kaon as the highest momentum particle in a hadronic jet. They also use a lifetime tag to independently identify a c jet, based on the impact parameter distribution of charged particles in a jet. From this measurement $|V_{CS}|$ is determined to be $0.94^{+0.32}_{-0.26}\pm0.13$.

AVERAGE PARTICLE MULTIPLICITIES IN HADRONIC W DECAY

Summed over particle and antiparticle, when appropriate.

 $\langle N_{\pi^{\pm}} \rangle$ $\frac{DOCUMENT\ ID}{49}$ ABREU,P 00F DLPH $\frac{ee}{cm} = 189\ GeV$ 15.70 ± 0.35

 $\langle N_{K^{\pm}} \rangle$ $\frac{DOCUMENT\ ID}{50}$ ABREU,P 00F DLPH $E_{\mathrm{cm}}^{ee}=189\ \mathrm{GeV}$ **VALUE** 2.20 ± 0.19

 $^{^{49}}$ ABREU,P 00F measure $\langle \mathit{N}_{\pi^\pm}
angle = 31.65 \pm 0.48 \pm 0.76$ and $15.51 \pm 0.38 \pm 0.40$ in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

 $^{^{50}}$ ABREU,P 00F measure $\langle \mathit{N}_{\mathit{K}^{\pm}} \rangle =$ 4.38 \pm 0.42 \pm 0.12 and 2.23 \pm 0.32 \pm 0.17 in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

<	Np	\rangle
•		,

 51 ABREU,P 00F measure $\langle N_p \rangle = 1.82 \pm 0.29 \pm 0.16$ and 0.94 \pm 0.23 \pm 0.06 in the fully hadronic and semileptonic final states respectively. The value quoted is a weighted average without assuming any correlations.

$\langle N_{\rm charged} \rangle$

VALUE	DOCUMENT ID		TECN	COMMENT
19.39±0.08 OUR AVERAGE				
$19.38\!\pm\!0.05\!\pm\!0.08$	⁵² ABBIENDI	06A	OPAL	E ^{ee} _{cm} = 189–209 GeV
19.44 ± 0.17	⁵³ ABREU,P	00F	DLPH	$E_{cm}^{ee} = 183 + 189 \text{ GeV}$
$19.3 \pm 0.3 \pm 0.3$	⁵⁴ ABBIENDI	99N	OPAL	$E_{\rm cm}^{\it ee}=183~{\rm GeV}$
19.23 ± 0.74	⁵⁵ ABREU	98C	DLPH	$E_{\rm cm}^{\rm ee} = 172 \; {\rm GeV}$

- 52 ABBIENDI 06A measure $\langle N_{\rm charged} \rangle = 38.74 \pm 0.12 \pm 0.26$ when both W bosons decay hadronically and $\langle N_{\rm charged} \rangle = 19.39 \pm 0.11 \pm 0.09$ when one W boson decays semileptonically. The value quoted here is obtained under the assumption that there is no color reconnection between W bosons; the value is a weighted average taking into account correlations in the systematic uncertainties.
- 53 ABREU,P 00F measure $\left< N_{\rm charged} \right> = 39.12 \pm 0.33 \pm 0.36$ and $38.11 \pm 0.57 \pm 0.44$ in the fully hadronic final states at 189 and 183 GeV respectively, and $\left< N_{\rm charged} \right> = 19.49 \pm 0.31 \pm 0.27$ and $19.78 \pm 0.49 \pm 0.43$ in the semileptonic final states. The value quoted is a weighted average without assuming any correlations.
- 54 ABBIENDI 99N use the final states $W^+W^- o q \overline{q} \ell \overline{\nu}_\ell$ to derive this value.
- 55 ABREU 98C combine results from both the fully hadronic as well semileptonic WW final states after demonstrating that the W decay charged multiplicity is independent of the topology within errors.

TRIPLE GAUGE COUPLINGS (TGC'S)

A REVIEW GOES HERE - Check our WWW List of Reviews



OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at http://lepewwg.web.cern.ch/LEPEWWG/lepww/tgc).

VALUE	<u>EVTS</u>	DOCUMENT ID		TECN	COMMENT
$0.984^{+0.022}_{-0.019}$ OUR FIT					
$1.001 \pm 0.027 \pm 0.013$	9310	⁵⁶ SCHAEL	05A	ALEP	E ^{ee} _{cm} = 183–209 GeV
$0.987 ^{+ 0.034}_{- 0.033}$	9800	⁵⁷ ABBIENDI	04 D	OPAL	E ^{ee} _{cm} = 183–209 GeV
$0.966^{+0.034}_{-0.032}{\pm0.015}$	8325	⁵⁸ ACHARD	04 D	L3	Eee = 161–209 GeV
• • • We do not use th	e followir	ng data for averages	s, fits,	limits, e	etc. • • •
	2.3	⁵⁹ ABAZOV	05 S	D0	$E_{cm}^{ar{p}} = 1.96 \; TeV$
$0.98 \ \pm 0.07 \ \pm 0.01$	2114	⁶⁰ ABREU	011	DLPH	E ^{ee} _{cm} = 183+189 GeV
	331	⁶¹ ABBOTT	991	D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$

- 56 SCHAEL 05A study single–photon, single–W, and WW–pair production from 183 to 209 GeV. The result quoted here is derived from the WW–pair production sample. Each parameter is determined from a single–parameter fit in which the other parameters assume their Standard Model values.
- ⁵⁷ ABBIENDI 04D combine results from W^+W^- in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $0.923 < g_1^Z < 1.054$.
- 58 ACHARD 04D study WW—pair production, single—W production and single—photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained from the WW—pair production sample including data from 161 to 183 GeV, ACCIA-RRI 99Q. Each parameter is determined from a single—parameter fit in which the other parameters assume their Standard Model values.
- 59 ABAZOV 05S study $\overline{p}\, p \to WZ$ production with a subsequent trilepton decay to $\ell\nu\ell'\overline{\ell}'$ (ℓ and $\ell'=e$ or μ). Three events (estimated background 0.71 \pm 0.08 events) with WZ decay characteristics are observed from which they derive limits on the anomalous WWZ couplings. The 95% CL limit for a form factor scale $\Lambda=1.5$ TeV is 0.51 < g_1^Z < 1.66, fixing λ_Z and κ_Z to their Standard Model values.
- 60 ABREU 011 combine results from e^+e^- interactions at 189 GeV leading to W^+W^- and $W\,e\,\nu_e$ final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is 0.84 $< g_1^Z < 1.13$.
- 61 ABBOTT 99I perform a simultaneous fit to the $W\gamma,~WW\to~$ dilepton, $WW/WZ\to e\nu jj,~WW/WZ\to~\mu\nu jj,~$ and $WZ\to~$ trilepton data samples. For $\Lambda=2.0$ TeV, the 95%CL limits are $0.63< g_1^Z<1.57,$ fixing λ_Z and κ_Z to their Standard Model values, and assuming Standard Model values for the $WW\gamma$ couplings.

κ_{γ}

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at http://lepewwg.web.cern.ch/LEPEWWG/lepww/tgc).

VALUE	<u>EVTS</u>	DOCUMENT ID		TECN	COMMENT	_
0.973 ^{+0.044} _{-0.045} OUR FIT	•					
$0.971\!\pm\!0.055\!\pm\!0.030$	10689	⁶² SCHAEL	05A	ALEP	E ^{ee} _{cm} = 183–209 GeV	
$0.88 \begin{array}{l} +0.09 \\ -0.08 \end{array}$	9800	⁶³ ABBIENDI	04 D	OPAL	E ^{ee} _{cm} = 183–209 GeV	
$1.013 { + 0.067 \atop - 0.064 } \pm 0.026$	10575	⁶⁴ ACHARD	04 D	L3	Eee = 161–209 GeV	
• • • We do not use the	ne followir	ng data for averages	s, fits,	limits, e	etc. • • •	
	17	⁶⁵ ABAZOV	06н	D0	$E_{cm}^{oldsymbol{p}\overline{oldsymbol{p}}}=1.96\;TeV$	
	141	⁶⁶ ABAZOV	05 J	D0	$E_{cm}^{p\overline{p}} = 1.96 \; TeV$	
$1.25 \ ^{+0.21}_{-0.20} \ \pm 0.06$	2298	⁶⁷ ABREU	011	DLPH	E ^{ee} _{cm} = 183+189 GeV	
		⁶⁸ BREITWEG	00	ZEUS	$e^+ p \rightarrow e^+ W^{\pm} X$, $\sqrt{s} \approx 300 \text{ GeV}$	
0.92 ± 0.34	331	⁶⁹ ABBOTT	991	D0	$\sqrt{s} \approx 300 \text{ GeV}$ $E_{\text{cm}}^{p\overline{p}} = 1.8 \text{ TeV}$	

- 62 SCHAEL 05A study single–photon, single–W, and WW–pair production from 183 to 209 GeV. Each parameter is determined from a single–parameter fit in which the other parameters assume their Standard Model values.
- 63 ABBIENDI 04D combine results from $W^+\,W^-$ in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $0.73 < \kappa_{\gamma} < 1.07$.
- 64 ACHARD 04D study WW-pair production, single-W production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 65 ABAZOV 06H study $\overline{p}p \to WW$ production with a subsequent decay $WW \to e^+\nu_e\,e^-\overline{\nu}_e,\,WW \to e^\pm\nu_e\,\mu^\mp\nu_\mu$ or $WW \to \mu^+\nu_\mu\mu^-\overline{\nu}_\mu$. The 95% C.L. limit for a form factor scale $\Lambda=1$ TeV is $-0.05<\kappa_\gamma<2.29$, fixing $\lambda_\gamma=0$. With the assumption that the $WW\gamma$ and WWZ couplings are equal the 95% C.L. one-dimensional limit ($\Lambda=2$ TeV) is $0.68<\kappa<1.45$.
- 66 ABAZOV 05J perform a likelihood fit to the photon E_T spectrum of $W\gamma+{\rm X}$ events, where the W decays to an electron or muon which is required to be well separated from the photon. For $\Lambda=2.0$ TeV the 95% CL limits are 0.12 < κ_{γ} < 1.96. In the fit λ_{γ} is kept fixed to its Standard Model value.
- 67 ABREU 011 combine results from $e^+\,e^-$ interactions at 189 GeV leading to $W^+\,W^-$, $W\,e\,\nu_e$, and $\nu\overline{\nu}\gamma$ final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is 0.87 $<\kappa_\gamma<1.68$.
- 68 BREITWEG 00 search for W production in events with large hadronic p_T . For $p_T>$ 20 GeV, the upper limit on the cross section gives the 95%CL limit $-3.7<\kappa_{\gamma}<$ 2.5 (for $\lambda_{\gamma}=$ 0).
- 69 ABBOTT 99I perform a simultaneous fit to the $W\gamma$, $WW\to dilepton$, $WW/WZ\to e\nu jj$, $WW/WZ\to \mu\nu jj$, and $WZ\to trilepton data samples. For <math>\Lambda=2.0$ TeV, the 95%CL limits are $0.75<\kappa_{\gamma}<1.39$.

 λ_{γ}

OUR FIT below is obtained by combining the measurements taking into account properly the common systematic errors (see LEPEWWG/TGC/2005-01 at $\frac{1}{LEPEWWG/lepeww}$).

VALUE	<u>EVTS</u>	DOCUMENT ID		TECN	COMMENT
$-0.028^{igoplus 0.020}_{-0.021}$ OUR F	IT				
$-0.012\!\pm\!0.027\!\pm\!0.011$	10689	⁷⁰ SCHAEL	05A	ALEP	$E_{\rm cm}^{\it ee} = 183 – 209 \; {\rm GeV}$
$-0.060^{+0.034}_{-0.033}$	9800	⁷¹ ABBIENDI	04 D	OPAL	E ^{ee} _{cm} = 183–209 GeV
$-0.021^{+0.035}_{-0.034}{\pm0.017}$	10575	⁷² ACHARD	04 D	L3	E ^{ee} _{cm} = 161–209 GeV
\bullet \bullet We do not use the	ne followir	ng data for averages	s, fits,	limits, e	etc. • • •
	17	⁷³ ABAZOV	06н	D0	$E_{Cm}^{p\overline{p}}=1.96\;TeV$
	141	⁷⁴ ABAZOV	05 J	D0	$E_{cm}^{p\overline{p}} = 1.96 \; TeV$
$0.05\ \pm0.09\ \pm0.01$	2298	⁷⁵ ABREU	011	DLPH	$E_{\rm cm}^{\it ee} = 183 + 189 \; {\rm GeV}$
		⁷⁶ BREITWEG	00		$e^+ p \rightarrow e^+ W^{\pm} X$, $\sqrt{s} \approx 300 \text{ GeV}$
$0.00 \begin{array}{l} +0.10 \\ -0.09 \end{array}$	331	77 ABBOTT	991	D0	$E_{cm}^{p\overline{p}} = 1.8 \; TeV$

- 70 SCHAEL 05A study single–photon, single–W, and WW–pair production from 183 to 209 GeV. Each parameter is determined from a single–parameter fit in which the other parameters assume their Standard Model values.
- 71 ABBIENDI 04D combine results from $W^+\,W^-$ in all decay channels. Only *CP*-conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is $-0.13 < \lambda_{\gamma} < 0.01$.
- 72 ACHARD 04D study WW-pair production, single-W production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained including data from 161 to 183 GeV, ACCIARRI 99Q. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 73 ABAZOV 06H study $\overline{p}p \to WW$ production with a subsequent decay $WW \to e^+\nu_e\,e^-\overline{\nu}_e$, $WW \to e^\pm\nu_e\,\mu^\mp\nu_\mu$ or $WW \to \mu^+\nu_\mu\,\mu^-\overline{\nu}_\mu$. The 95% C.L. limit for a form factor scale $\Lambda=1$ TeV is $-0.97 < \lambda_\gamma < 1.04$, fixing $\kappa_\gamma=1$. With the assumption that the $WW\gamma$ and WWZ couplings are equal the 95% C.L. one-dimensional limit ($\Lambda=2$ TeV) is $-0.29 < \lambda < 0.30$.
- 74 ABAZOV 05J perform a likelihood fit to the photon E_T spectrum of $W\gamma+{\rm X}$ events, where the W decays to an electron or muon which is required to be well separated from the photon. For $\Lambda=2.0$ TeV the 95% CL limits are $-0.20<\lambda_{\gamma}<0.20$. In the fit κ_{γ} is kept fixed to its Standard Model value.
- 75 ABREU 011 combine results from e^+e^- interactions at 189 GeV leading to W^+W^- , $We\nu_e$, and $\nu\overline{\nu}\gamma$ final states with results from ABREU 99L at 183 GeV. The 95% confidence interval is $-0.11<\lambda_{\gamma}<0.23$.
- 76 BREITWEG 00 search for W production in events with large hadronic p_T . For $p_T>$ 20 GeV, the upper limit on the cross section gives the 95%CL limit $-3.2<\lambda_{\gamma}<3.2$ for κ_{γ} fixed to its Standard Model value.
- 77 ABBOTT 99I perform a simultaneous fit to the $W\gamma,~WW\to$ dilepton, $WW/WZ\to e\nu jj,~WW/WZ\to \mu\nu jj,$ and $WZ\to$ trilepton data samples. For $\Lambda=2.0$ TeV, the 95%CL limits are $-0.18<\lambda_{\gamma}<0.19.$

κ_Z

This coupling is *CP*-conserving (*C*- and *P*- separately conserving).

<u>VALUE</u>	<u>EVTS</u>	DOCUMENT ID		<u>TECN</u>	COMMENT
$0.924^{+0.059}_{-0.056}\pm0.024$	7171	⁷⁸ ACHARD	04 D	L3	$E_{\rm cm}^{\it ee} = 189 – 209 \; {\rm GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

17
79
 ABAZOV 06H D0 $E_{\rm cm}^{p\overline{p}}=1.96$ TeV 2.3 80 ABAZOV 05S D0 $E_{\rm cm}^{p\overline{p}}=1.96$ TeV

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- 78 ACHARD 04D study WW-pair production, single-W production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the WW-pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.
- 79 ABAZOV 06H study $\overline{p}p \to WW$ production with a subsequent decay $WW \to e^+\nu_e\,e^-\overline{\nu}_e$, $WW \to e^\pm\nu_e\,\mu^\mp\nu_\mu$ or $WW \to \mu^+\nu_\mu\,\mu^-\overline{\nu}_\mu$. The 95% C.L. limit for a form factor scale $\Lambda=2$ TeV is 0.55 $<\kappa_Z<$ 1.55, fixing $\lambda_Z=$ 0. With the assumption that the $WW\gamma$ and WWZ couplings are equal the 95% C.L. one-dimensional limit ($\Lambda=2$ TeV) is 0.68 $<\kappa<1.45$.
- ⁸⁰ ABAZOV 05S study $\overline{p}\,p \to WZ$ production with a subsequent trilepton decay to $\ell\nu\ell'\overline{\ell}'$ (ℓ and $\ell'=e$ or μ). Three events (estimated background 0.71 ± 0.08 events) with WZ decay characteristics are observed from which they derive limits on the anomalous WWZ couplings. The 95% CL limit for a form factor scale $\Lambda=1$ TeV is $-1.0<\kappa_Z<3.4$, fixing λ_Z and g_1^Z to their Standard Model values.

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 λ_{Z}

This coupling is *CP*-conserving (*C*- and *P*- separately conserving).

<u>VALUE</u>	<u>EVTS</u>	DOCUMENT ID		-	COMMENT
$-0.088^{+0.060}_{-0.057}\pm0.023$	7171	⁸¹ ACHARD	04 D	L3	$E_{\sf cm}^{\it ee} = 189 – 209 \; {\sf GeV}$

• • We do not use the following data for averages, fits, limits, etc.

17 82 ABAZOV 06H D0
$$E_{\rm cm}^{p\overline{p}}=1.96~{\rm TeV}$$

2.3 83 ABAZOV 05S D0 $E_{\rm cm}^{p\overline{p}}=1.96~{\rm TeV}$

 81 ACHARD 04D study WW-pair production, single-W production and single-photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the WW-pair production sample. Each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values.

⁸² ABAZOV 06H study $\overline{p}p \to WW$ production with a subsequent decay $WW \to e^+\nu_e\,e^-\overline{\nu}_e$, $WW \to e^\pm\nu_e\,\mu^\mp\nu_\mu$ or $WW \to \mu^+\nu_\mu\,\mu^-\overline{\nu}_\mu$. The 95% C.L. limit for a form factor scale $\Lambda=2$ TeV is $-0.39 < \lambda_Z < 0.39$, fixing $\kappa_Z=1$. With the assumption that the $WW\gamma$ and WWZ couplings are equal the 95% C.L. one-dimensional limit ($\Lambda=2$ TeV) is $-0.29 < \lambda < 0.30$.

⁸³ ABAZOV 05S study $\overline{p}\,p \to WZ$ production with a subsequent trilepton decay to $\ell\nu\ell'\overline{\ell}'$ (ℓ and $\ell'=e$ or μ). Three events (estimated background 0.71 \pm 0.08 events) with WZ decay characteristics are observed from which they derive limits on the anomalous WWZ couplings. The 95% CL limit for a form factor scale $\Lambda=1.5$ TeV is $-0.48<\lambda_Z<0.48$, fixing g_1^Z and κ_Z to their Standard Model values.

 g_5^Z

This coupling is CP-conserving but C- and P-violating.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT				
0.93±0.09 OUR AVERAGE Error includes scale factor of 1.1.								
$0.96^{+0.13}_{-0.12}$	9800	⁸⁴ ABBIENDI	04D OPAL	<i>E</i> ^{ee} _{cm} = 183−209 GeV				
$1.00\!\pm\!0.13\!\pm\!0.05$	7171	⁸⁵ ACHARD	04D L3	E ^{ee} _{cm} = 189–209 GeV				
$0.56^{igoplus 0.23}_{-0.22} \!\pm\! 0.12$	1154	⁸⁶ ACCIARRI	99Q L3	Eee = 161+172+ 183 GeV				

• • • We do not use the following data for averages, fits, limits, etc. • • •

 0.84 ± 0.23 87 EBOLI 00 THEO LEP1, SLC+ Tevatron

 84 ABBIENDI 04D combine results from $W^+\,W^-$ in all decay channels. Only $\it CP$ -conserving couplings are considered and each parameter is determined from a single-parameter fit in which the other parameters assume their Standard Model values. The 95% confidence interval is 0.72 $<\!g_5^Z<1.21.$

- 85 ACHARD 04D study WW—pair production, single—W production and single—photon production with missing energy from 189 to 209 GeV. The result quoted here is obtained using the WW—pair production sample. Each parameter is determined from a single—parameter fit in which the other parameters assume their Standard Model values.
- 86 ACCIARRI 99Q study W-pair, single-W, and single photon events.
- ⁸⁷ EBOLI 00 extract this indirect value of the coupling studying the non-universal one-loop contributions to the experimental value of the $Z \rightarrow b \bar{b}$ width (Λ =1 TeV is assumed).

 g_{Δ}^{Z}

This coupling is CP-violating (C-violating and P-conserving).

coap 6 .c	0	6 (55		<i>)</i> ·	
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
$-0.02 { +0.32 \atop -0.33 }$	1065	⁸⁸ ABBIENDI	01н OPAL	<i>E</i> ^{ee} _{cm} = 189 GeV	

⁸⁸ ABBIENDI 01H study W-pair events, with one leptonically and one hadronically decaying W. The coupling is extracted using information from the W production angle together with decay angles from the leptonically decaying W.

$\widetilde{\kappa}_{Z}$

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

VALUE		EVTS	DOCUME	ENT ID	TECN_	COMMENT
$-0.20^{+0.10}_{-0.07}$		1065	⁸⁹ ABBIEN	NDI 01ı	H OPAL	E ^{ee} _{cm} = 189 GeV

 $^{^{89}}$ ABBIENDI 01H study W-pair events, with one leptonically and one hadronically decaying W. The coupling is extracted using information from the W production angle together with decay angles from the leptonically decaying W.

$\widetilde{\lambda}_{\pmb{Z}}$

This coupling is *CP*-violating (*C*-conserving and *P*-violating).

<u>VALUE</u>	•	 <u>EVTS</u>	DOCUMENT ID		<u>TECN</u>	COMMENT	
$-0.18^{+0.24}_{-0.16}$		1065	⁹⁰ ABBIENDI	01H	OPAL	<i>E</i> ^{ee} _{cm} = 189 GeV	

 $^{^{90}}$ ABBIENDI 01H study W-pair events, with one leptonically and one hadronically decaying W. The coupling is extracted using information from the W production angle together with decay angles from the leptonically decaying W.

W ANOMALOUS MAGNETIC MOMENT

The full magnetic moment is given by $\mu_W=e(1+\kappa+\lambda)/2m_W$. In the Standard Model, at tree level, $\kappa=1$ and $\lambda=0$. Some papers have defined $\Delta\kappa=1-\kappa$ and assume that $\lambda=0$. Note that the electric quadrupole moment is given by $-e(\kappa-\lambda)/m_W^2$. A description of the parameterization of these moments and additional references can be found in HAGIWARA 87 and BAUR 88. The parameter Λ appearing in the theoretical limits below is a regularization cutoff which roughly corresponds to the energy scale where the structure of the W boson becomes manifest.

VALUE (e/2m _W)	EVTS	DOCUMENT ID		TECN	COMMENT
2.22 ^{+0.20} -0.19	2298	⁹¹ ABREU	011	DLPH	E ^{ee} _{cm} = 183+189 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

⁹² ABE	95 G	CDF
⁹³ ALITTI	92C	UA2
⁹⁴ SAMUEL	92	THEO
⁹⁵ SAMUEL	91	THEO
⁹⁶ GRIFOLS	88	THEO
⁹⁷ GROTCH	87	THEO
⁹⁸ VANDERBIJ	87	THEO
⁹⁹ GRAU	85	THEO
¹⁰⁰ SUZUKI	85	THEO
¹⁰¹ HERZOG	84	THEO

- ⁹¹ ABREU 011 combine results from e^+e^- interactions at 189 GeV leading to W^+W^- , $We\nu_e$, and $\nu\overline{\nu}\gamma$ final states with results from ABREU 99L at 183 GeV to determine Δg_1^Z , $\Delta \kappa_\gamma$, and λ_γ . $\Delta \kappa_\gamma$ and λ_γ are simultaneously floated in the fit to determine μ_W .
- 92 ABE 95G report $-1.3 < \kappa < 3.2$ for $\lambda=0$ and $-0.7 < \lambda < 0.7$ for $\kappa=1$ in $p\overline{p} \to e \nu_e \gamma X$ and $\mu \nu_\mu \gamma X$ at $\sqrt{s}=1.8$ TeV.
- 93 ALITTI 92C measure $\kappa=1^{+2.6}_{-2.2}$ and $\lambda=0^{+1.7}_{-1.8}$ in $p\overline{p}\to e\nu\gamma+$ X at $\sqrt{s}=630$ GeV. At 95%CL they report $-3.5<\kappa<5.9$ and $-3.6<\lambda<3.5$.
- $^{94}\, {\sf SAMUEL}$ 92 use preliminary CDF and UA2 data and find $-2.4 < \kappa < 3.7$ at 96%CL and $-3.1 < \kappa < 4.2$ at 95%CL respectively. They use data for $W\,\gamma$ production and radiative W decay.
- ⁹⁵ SAMUEL 91 use preliminary CDF data for $p\overline{p}\to W\gamma X$ to obtain $-11.3 \le \Delta\kappa \le 10.9$. Note that their $\kappa=1-\Delta\kappa$.
- ⁹⁶ GRIFOLS 88 uses deviation from ρ parameter to set limit $\Delta \kappa \lesssim 65 \ (M_W^2/\Lambda^2)$.
- 97 GROTCH 87 finds the limit $-37 < \Delta \kappa < 73.5$ (90% CL) from the experimental limits on $e^+\,e^- \to \, \nu \overline{\nu} \gamma$ assuming three neutrino generations and $-19.5 < \Delta \kappa < 56$ for four generations. Note their $\Delta \kappa$ has the opposite sign as our definition.
- ⁹⁸ VANDERBIJ 87 uses existing limits to the photon structure to obtain $|\Delta\kappa| < 33$ (m_W/Λ) . In addition VANDERBIJ 87 discusses problems with using the ρ parameter of the Standard Model to determine $\Delta\kappa$.
- GRAU 85 uses the muon anomaly to derive a coupled limit on the anomalous magnetic dipole and electric quadrupole (λ) moments $1.05 > \Delta \kappa \ln(\Lambda/m_W) + \lambda/2 > -2.77$. In the Standard Model $\lambda = 0$.
- SUZUKI 85 uses partial-wave unitarity at high energies to obtain $|\Delta\kappa|\lesssim 190~(m_W/\Lambda)^2$. From the anomalous magnetic moment of the muon, SUZUKI 85 obtains $|\Delta\kappa|\lesssim 2.2/\ln(\Lambda/m_W)$. Finally SUZUKI 85 uses deviations from the ρ parameter and obtains a very qualitative, order-of-magnitude limit $|\Delta\kappa|\lesssim 150~(m_W/\Lambda)^4$ if $|\Delta\kappa|\ll 1$.
- 101 HERZOG 84 consider the contribution of W-boson to muon magnetic moment including anomalous coupling of $WW\gamma$. Obtain a limit $-1 < \Delta\kappa < 3$ for $\Lambda \gtrsim 1$ TeV.

ANOMALOUS W/Z QUARTIC COUPLINGS

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a_0/Λ^2 , a_c/Λ^2 , a_n/Λ^2

Using the $WW\gamma$ final state, the LEP combined 95% CL limits on the anomalous contributions to the $WW\gamma\gamma$ and $WWZ\gamma$ vertices (as of summer 2003) are given below:

(See P. Wells, "Experimental Tests of the Standard Model," Int. Europhysics Conference on High-Energy Physics, Aachen, Germany, 17–23 July 2003)

$$-0.02 < a_0^W/\Lambda^2 < 0.02 \text{ GeV}^{-2},$$

 $-0.05 < a_c^W/\Lambda^2 < 0.03 \text{ GeV}^{-2},$
 $-0.15 < a_n/\Lambda^2 < 0.15 \text{ GeV}^{-2}.$

VALUE <u>DOCUMENT ID</u> <u>TECN</u>

• • • We do not use the following data for averages, fits, limits, etc. • • •

102 ABBIENDI04B OPAL103 ABBIENDI04L OPAL104 HEISTER04A ALEP105 ABDALLAH03I DLPH106 ACHARD02F L3

- 102 ABBIENDI 04B select 187 $e^+e^-\to W^+W^-\gamma$ events in the C.M. energy range 180–209 GeV, where $E_{\gamma}>$ 2.5 GeV, the photon has a polar angle $|\cos\!\theta_{\gamma}|<$ 0.975 and is well isolated from the nearest jet and charged lepton, and the effective masses of both fermion-antifermion systems agree with the W mass within 3 Γ_W . The measured differential cross section as a function of the photon energy and photon polar angle is used to extract the 95% CL limits: $-0.020~{\rm GeV}^{-2} < a_0/\Lambda^2 < 0.020~{\rm GeV}^{-2},$ $-0.053~{\rm GeV}^{-2} < a_c/\Lambda^2 < 0.037~{\rm GeV}^{-2}$ and $-0.16~{\rm GeV}^{-2} < a_n/\Lambda^2 < 0.15~{\rm GeV}^{-2}$.
- ABBIENDI 04L select 20 $e^+e^- ou \overline{
 u}\gamma\gamma$ acoplanar events in the energy range 180–209 GeV and 176 $e^+e^- ou q\overline{q}\gamma\gamma$ events in the energy range 130–209 GeV. These samples are used to constrain possible anomalous $W^+W^-\gamma\gamma$ and $ZZ\gamma\gamma$ quartic couplings. Further combining with the $W^+W^-\gamma$ sample of ABBIENDI 04B the following one-parameter 95% CL limits are obtained: $-0.007 < a_0^Z/\Lambda^2 < 0.023 \ {\rm GeV}^{-2}, -0.029 < a_c^Z/\Lambda^2 < 0.029 \ {\rm GeV}^{-2}, -0.020 < a_0^W/\Lambda^2 < 0.020 \ {\rm GeV}^{-2}, -0.052 < a_c^W/\Lambda^2 < 0.037 \ {\rm GeV}^{-2}$
- In the CM energy range 183 to 209 GeV HEISTER 04A select 30 $e^+\,e^-\to\nu\overline{\nu}\gamma\gamma$ events with two acoplanar, high energy and high transverse momentum photons. The photon-photon acoplanarity is required to be > 5°, $E_{\gamma}/\sqrt{s}>$ 0.025 (the more energetic photon having energy > 0.2 \sqrt{s}), p $_{T_{\gamma}}/\rm E_{beam}>$ 0.05 and $|\cos\theta_{\gamma}|<$ 0.94. A likelihood fit to the photon energy and recoil missing mass yields the following one–parameter 95% CL limits: $-0.012< a_0^Z/\Lambda^2<0.019~{\rm GeV}^{-2}, -0.041< a_c^Z/\Lambda^2<0.044~{\rm GeV}^{-2}, -0.060< a_0^W/\Lambda^2<0.055~{\rm GeV}^{-2}, -0.099< a_c^W/\Lambda^2<0.093~{\rm GeV}^{-2}.$
- ABDALLAH 03I select 122 e⁺e⁻ \rightarrow W⁺W⁻ γ events in the C.M. energy range 189–209 GeV, where $E_{\gamma} >$ 5 GeV, the photon has a polar angle $|\cos\theta_{\gamma}| <$ 0.95 and is well isolated from the nearest charged fermion. A fit to the photon energy spectra yields $a_c/\Lambda^2 = 0.000^{+0.019}_{-0.040} \; {\rm GeV^{-2}}, \; a_0/\Lambda^2 = -0.004^{+0.018}_{-0.010} \; {\rm GeV^{-2}}, \; \widetilde{a_0}/\Lambda^2 = -0.007^{+0.019}_{-0.008} \; {\rm GeV^{-2}}, \; a_n/\Lambda^2 = -0.09^{+0.16}_{-0.05} \; {\rm GeV^{-2}}, \; {\rm and} \; \widetilde{a_n}/\Lambda^2 = +0.05^{+0.07}_{-0.15} \; {\rm GeV^{-2}}, \; {\rm keeping} \; {\rm the} \; {\rm other} \; {\rm parameters} \; {\rm fixed} \; {\rm to} \; {\rm their} \; {\rm Standard} \; {\rm Model} \; {\rm values} \; (0). \; {\rm The} \; 95\% \; {\rm CL} \; {\rm limits} \; {\rm are} : \; -0.063 \; {\rm GeV^{-2}} \; < a_c/\Lambda^2 \; < \; +0.032 \; {\rm GeV^{-2}}, \; -0.020 \; {\rm GeV^{-2}} \; < a_0/\Lambda^2 \; < \; +0.020 \; {\rm GeV^{-2}}, \; -0.020 \; {\rm GeV^{-2}} \; < a_n/\Lambda^2 \; < \; +0.020 \; {\rm GeV^{-2}}, \; -0.18 \; {\rm GeV^{-2}} \; < a_n/\Lambda^2 \; < \; +0.14 \; {\rm GeV^{-2}}, \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.16 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.18 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.18 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.18 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.18 \; {\rm GeV^{-2}} \; < \widetilde{a_n}/\Lambda^2 \; < \; +0.17 \; {\rm GeV^{-2}}. \; -0.18 \; {\rm Ge$
- 4CHARD 02F select 86 e⁺e[−] → W⁺W[−]γ events at 192–207 GeV, where $E_{\gamma} > 5$ GeV and the photon is well isolated. They also select 43 acoplanar e⁺e[−] → $\nu \overline{\nu} \gamma \gamma$ events in this energy range, where the photon energies are >5 GeV and >1 GeV and the photon polar angles are between 14° and 166°. All these 43 events are in the recoil mass region corresponding to the Z (75–110 GeV). Using the shape and normalization of the photon spectra in the $W^+W^-\gamma$ events, and combining with the 42 event sample from 189 GeV data (ACCIARRI 00T), they obtain: $a_0/\Lambda^2 = 0.000 \pm 0.010$ GeV^{−2}, $a_c/\Lambda^2 = -0.013 \pm 0.023$ GeV^{−2}, and $a_n/\Lambda^2 = -0.002 \pm 0.076$ GeV^{−2}. Further combining the analyses of $W^+W^-\gamma$ events with the low recoil mass region of $\nu \overline{\nu} \gamma \gamma$ events (including samples collected at 183 + 189 GeV), they obtain the following one-parameter 95% CL limits: -0.015 GeV^{−2} $< a_0/\Lambda^2 < 0.015$ GeV^{−2}, -0.048 GeV^{−2} $< a_c/\Lambda^2 < 0.026$ GeV^{−2}, and -0.14 GeV^{−2} $< a_n/\Lambda^2 < 0.13$ GeV^{−2}.

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EBOLI ABBIENDI ABBOTT ABBOTT ABREU ACCIARRI ACCIARRI BARATE BARATE BARATE ABBOTT ABBOTT	00D 00 99N 99H 99I 99L 99 99Q 99I 99L 99M 98N 98P	EPJ C12 411 MPL A15 1 PL B453 153 PR D60 052003 PR D60 072002 PL B459 382 PL B454 386 PL B467 171 PL B453 107 PL B462 389 PL B465 349 PR D58 092003 PR D58 012002	J. Breitweg et al. O. Eboli, M. Gonzalez-Garcia, S. G. Abbiendi et al. B. Abbott et al. B. Abreu et al. P. Abreu et al. M. Acciarri et al. M. Acciarri et al. R. Barate et al. R. Barate et al. R. Barate et al. B. Abbott et al. B. Abbott et al. B. Abbott et al.	(ZEUS Collab.) Novaes (OPAL Collab.) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.) (ALEPH Collab.) (ALEPH Collab.) (ALEPH Collab.) (ALEPH Collab.) (D0 Collab.) (D0 Collab.)
EBOLI ABBIENDI ABBOTT ABREU ACCIARRI ACCIARRI BARATE BARATE BARATE ABBOTT ABBOTT ABE	00D 00 99N 99H 99I 99L 99 99Q 99I 99L 99M 98N 98P 98H	EPJ C12 411 MPL A15 1 PL B453 153 PR D60 052003 PR D60 072002 PL B459 382 PL B454 386 PL B467 171 PL B453 107 PL B462 389 PL B465 349 PR D58 092003 PR D58 012002 PR D58 031101	J. Breitweg et al. O. Eboli, M. Gonzalez-Garcia, S. G. Abbiendi et al. B. Abbott et al. B. Abreu et al. P. Abreu et al. M. Acciarri et al. M. Acciarri et al. R. Barate et al. R. Barate et al. B. Abbott et al. F. Abe et al.	(ZEUS Collab.) Novaes (OPAL Collab.) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.) (ALEPH Collab.) (ALEPH Collab.) (ALEPH Collab.) (ALEPH Collab.) (D0 Collab.) (CDF Collab.)
EBOLI ABBIENDI ABBOTT ABBOTT ABREU ACCIARRI ACCIARRI BARATE BARATE BARATE ABBOTT ABBOTT	00D 00 99N 99H 99I 99L 99 99Q 99I 99L 99M 98N 98P	EPJ C12 411 MPL A15 1 PL B453 153 PR D60 052003 PR D60 072002 PL B459 382 PL B454 386 PL B467 171 PL B453 107 PL B462 389 PL B465 349 PR D58 092003 PR D58 012002	J. Breitweg et al. O. Eboli, M. Gonzalez-Garcia, S. G. Abbiendi et al. B. Abbott et al. B. Abreu et al. P. Abreu et al. M. Acciarri et al. M. Acciarri et al. R. Barate et al. R. Barate et al. R. Barate et al. B. Abbott et al. B. Abbott et al. B. Abbott et al.	(ZEUS Collab.) Novaes (OPAL Collab.) (D0 Collab.) (D0 Collab.) (DELPHI Collab.) (L3 Collab.) (L3 Collab.) (ALEPH Collab.) (ALEPH Collab.) (ALEPH Collab.) (ALEPH Collab.) (D0 Collab.) (D0 Collab.)
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ALITTI	92D	PL B277 203	J. Alitti <i>et al.</i>	(UA2 Collab.)
ALITTI	92F	PL B280 137	J. Alitti <i>et al.</i>	(UA2 Collab.)
SAMUEL	92	PL B280 124	M.A. Samuel <i>et al.</i>	(OKSU, CARL)
ABE	91C	PR D44 29	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	91	PL B253 503	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALITTI	91C	ZPHY C52 209	J. Alitti et al.	(UA2 Collab.)
SAMUEL	91	PRL 67 9	M.A. Samuel et al.	(OKSU, CARL)
Also		PRL 67 2920 (erratum)	M.A. Samuel et al.	· · ·
ABE	90	PRL 64 152	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PR D44 29	F. Abe <i>et al.</i>	(CDF Collab.)
ABE	90G	PRL 65 2243	F. Abe <i>et al.</i>	(CDF Collab.)
Also		PR D43 2070	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	90	PL B241 283	C. Albajar <i>et al.</i>	(UA1 Collab.)
ALITTI	90B	PL B241 150	J. Alitti et al.	(UA2 Collab.)
ALITTI	90C	ZPHY C47 11	J. Alitti <i>et al.</i>	(UA2 Collab.)
ABE	89I	PRL 62 1005	F. Abe <i>et al.</i>	(CDF Collab.)
ALBAJAR	89	ZPHY C44 15	C. Albajar <i>et al.</i>	(UA1 Collab.)
BAUR	88	NP B308 127	U. Baur, D. Zeppenfeld	(FSU, WISC)
GRIFOLS	88	IJMP A3 225	J.A. Grifols, S. Peris, J. Sola	(BÀRC, DESY)
Also		PL B197 437	J.A. Grifols, S. Peris, J. Sola	(BARC, DESY)
ALBAJAR	87	PL B185 233	C. Albajar <i>et al.</i>	` (UA1 Collab.)
ANSARI	87	PL B186 440	R. Ansari et al.	(UA2 Collab.)
ANSARI	87C	PL B194 158	R. Ansari et al.	(UA2 Collab.)
GROTCH	87	PR D36 2153	H. Grotch, R.W. Robinett	` (PSU)
HAGIWARA	87	NP B282 253	K. Hagiwara et al.	(KEK, UCLA, FSU)
VANDERBIJ	87	PR D35 1088	J.J. van der Bij	` (FNAL)
APPEL	86	ZPHY C30 1	J.A. Appel <i>et al.</i>	(UA2 Čollab.)
ARNISON	86	PL 166B 484	G.T.J. Arnison et al.	(UA1 Collab.) J
ALTARELLI	85B	ZPHY C27 617	G. Altarelli, R.K. Ellis, G. Martinelli	` (CERN+)
GRAU	85	PL 154B 283	A. Grau, J.A. Grifols	(BARC)
SUZUKI	85	PL 153B 289	M. Suzuki	` (LBL)
ARNISON	84D	PL 134B 469	G.T.J. Arnison et al.	(UA1 Collab.)
HERZOG	84	PL 148B 355	F. Herzog	` (WISC)
Also		PL 155B 468 (erratum)	F. Herzog	(WISC)
ARNISON	83	PL 122B 103 `	G.T.J. Arnison et al.	(UA1 Čollab.)
BANNER	83B	PL 122B 476	M. Banner et al.	(UA2 Collab.)
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